M. Huszár

DEFORMATION OF THE SO $(2, C)$ SUBGROUP OF THE LORENTZ GROUP

Abungaxian Academy of sciences
CENTRAL RESEARCH
INSTITUTE FOR PHYSICS


BUDAPEST
$2017$

DEFORMATION OF THE SO ( $2, \mathrm{C}$ ) SUBGROUP OF THE LORENTZ GROUP
M. Huszár

Central Research Institute for Physics of the Hungarian Academy of Sciences, Budapest

## ABSTRACT

The two-dimensional complex sphere $s_{1}^{2}+s_{2}^{2}+s_{3}^{2}=s^{2}$ forms a homogeneous space under the $S L(2, C)$ group. The little group of a point in this space is the $S O(2, C)$ group or the horosheric group $T(2)$ according to whether $S \neq 0$ or $S=0$. Deformation of the $S O(2, C)$ group.into $T$ (2) is investigated and is demonstrated on unitary representations. This deformation is a counterpart to that of the little groups $S O(3), E(2), S O(2,1)$ belonging to the hyperboloid family.

## PE3iOME

Двумерная комплексная сфера $\mathrm{S}_{1}^{2}+\mathrm{S}_{2}^{2}+\mathrm{S}_{3}^{2}=\mathrm{S}^{2}$ образует однородное пространство относительно группы Лоренца. Малыми группами некоторой точки, находящейся в этом пространстве являются $S O(2, C)$ или группа орисферических трансляций $T(2)$ в зависимости от того, что $S \neq 0$ или $S=0$. В настоящей работе рассмотрена деформация группы $S O(2, C)$ в группу $T(2)$, а также продемонстрирована на соответствующих унитарных представлениях. Вышеуказанная деформация представляет собой аналогию деформаций друг в друга малых групп $S O(3), E(2), S O(2,1)$, принадлежащих к семейству гиперболоида.

## KIVONAT

$A z s_{1}^{2}+s_{2}^{2}+S_{3}^{2}=s^{2}$ egyenlet által leirt kétđimenziós komplex gömb homogén teret képez az SL (2,C) csoporttal szemben. Egy ebben a térben helyetfoglaló pont kiscsoportja az $S 0(2, C)$ illetve a $T(2)$ horoszférikús alcsoport, attól függóen, hogy $S \neq O$ vagy $S=0$. Jelen dolgozatban az $S O$ ( $2, C$ ) csoportnak a T (2) csoportba való deformációját vizsgáljuk és az unitér ábrázolásokon demonstráljuk. Ez a deformáció a hiperboloid családhoz tartozó SO (3), $E(2)$, SO $(2,1)$ kiscsoportok egymásba való deformálásának analogonja.
horospheric group isomorphic to the two-dimensional translation group $T(2)$. To this end a family of homogeneous spaces should be used the little groups of which are apt for demonstration of the deformation process. Since the proper Lorentz group is isomorphic to the connected part of the three-dimensional complex rotation group [5], the two-dimensional complex sphere

$$
s_{1}^{2}+s_{2}^{2}+s_{3}^{2}=s^{2}
$$

/hereafter $\Sigma_{S} /$ forms a homogeneous space under the proper Lorentz group, as well as under $S L(2, C)$ The three-dimensional complex vector $\underset{\sim}{S}=\left(S_{1}, S_{2}, S_{3}\right)$, which is the self-dual part of a Lorentz covariant antisymmetric tensor $S_{\mu \nu} / \mu, \nu=0,1,2,3 /$ under Ae $S L(2, C)$ transforms as follows

$$
\hat{S}^{\prime}=A \hat{S} A^{-1}
$$

where $\hat{S}=\underset{\sim}{\sigma} \underset{\sim}{S}=\sigma_{1} S_{1}+\sigma_{2} S_{2}+\sigma_{3} S_{3} \quad$ and $\quad \sigma_{i}-s \quad$ stand for the Pauli matrices. ${ }^{\text {T }}$ /

We choose a standard vector on $\Sigma_{0}$ as follows

$$
\underset{\sim}{S_{0}}=(0,0, S)
$$

Here $S$ is supposed to be non-zero. The little group of this vector, that is, the subgroup of $\mathrm{SL}(2, C)$ satisfying the condition

$$
\hat{\mathrm{S}}_{\mathrm{O}}=\mathrm{H}_{\mathrm{O}} \hat{\mathrm{~S}}_{\mathrm{O}} \mathrm{H}_{\mathrm{O}}^{-1}
$$

is clearly of the form

$$
H_{o}(\varphi)=e^{-i \frac{\varphi}{2} \sigma_{3}}=\left(\begin{array}{cc}
e^{-i \frac{\varphi}{2}} & 0 \\
0 & e^{i \frac{\varphi}{2}}
\end{array}\right)
$$

[^0]where $\varphi=\varphi_{1}+i \varphi_{2}$ is a complex angle with a real part describing a rotation about the $z$-axis and varying in the range $-2 \pi \leq \varphi_{1}<2 \pi$ and with an imaginary part describing a boost along the z-axis and varying in the range $-\infty<\varphi_{2}<\infty$. It follows that this group is $\mathrm{SO}(2, \mathrm{c})=\mathrm{So}(2) \times \mathrm{So}(1,1)$.

In a similar way, by choosing the standard vector

$$
S_{N \infty}=(S, \text { i } S, 0), \quad \mid S \neq 0 /
$$

on the complex sphere of zero radius $\Sigma_{0}$, we arrive at the horospheric little group isomorphic to the two-dimensional real Euclidean translation group $T(2)$,

$$
H_{\infty}(\varphi)=e^{-i \frac{\varphi}{2} \sigma_{+}}=\left(\begin{array}{cc}
1 & -i \varphi \\
0 & 1
\end{array}\right)
$$

where $\sigma_{+}=\sigma_{1}+i \sigma_{2}$ and $\varphi=\varphi_{1}+i \varphi_{2}$. In the present case both $\varphi_{1}$ and $\varphi_{2}$ vary from $-\infty$ to $\infty$. It is easy to see the validity of the inverse statement, that is, if the little group of a three-dimensional complex vector is $T(2) / \mathrm{SO}(2, \mathrm{c}) /$ then it is situated on the complex sphere of zero /non--zero/ radius. Here and throughout this paper the $\underset{\sim}{S}=0$ point is supposed to be excluded from $\Sigma_{o}$ since this point itself is invariant under $\mathrm{SL}(2, \mathrm{c})$. Hence by including $\underset{\sim}{S}=0$ the homogeneity of $\Sigma_{0}$ would be spoiled.

Consider now the vector

$$
{\underset{\sim}{S}}_{\tau}=\left(\frac{\tau}{1+\tau} S, \frac{i \tau}{1+\tau} S, \frac{1}{1+\tau} S\right)
$$

interpolating between $\underset{\sim}{S}{ }_{0}$ and $\underset{\sim}{S_{\infty}}$. Here $\tau$ is a real parameter describing the deformation varying in the range

$$
0 \leqslant \tau<\infty \quad .
$$

The limits of the vector $/ 1.8 /$ as $\tau \rightarrow 0$ and $\tau \rightarrow \infty$ are ${\underset{\sim}{\sim}}_{O}^{S}$ and ${\underset{\sim}{\infty}}_{\infty}$ as given by Eqs. $/ 1.3 /$ and /1.6/. Since the length of the vector ${\underset{\sim}{\sim}}^{S_{~}}$

$$
\left[\left(s_{\tau}\right)_{1}^{2}+\left(s_{\tau}\right)_{2}^{2}+\left(s_{\tau}\right)_{3}^{2}\right]^{1 / 2}=\frac{S^{1+\tau}}{11}
$$

is non-zero for $\tau<\infty$ the little group of ${\underset{\sim}{~}}_{\tau}$ is an $\operatorname{so}(2, c)$ group isomorphic to $H_{0}$. For ${ }_{\tau \rightarrow \infty}$ the little group $H_{\infty}$ as given by Eq. /1.7/ is obtained. By making use of Eqs. /1.2/ and /1.8/ we get an explicit form of the little group of ${\underset{\sim}{~}}_{\tau}$ for an arbitrary value of $\tau$ :

$$
H_{\tau}(\varphi)=\exp \left\{-i \frac{\varphi}{2}\left(\sigma_{3}+\tau \sigma_{+}\right) \frac{1}{1+\tau}\right\}=\left(\begin{array}{cc}
e^{-i \frac{1}{1+\tau} \frac{\varphi}{2}} & -2 i \tau \sin \left(\frac{1}{1+\tau} \frac{\varphi}{2}\right) \\
0 & e^{i \frac{1}{1+\tau} \frac{\varphi}{2}}
\end{array}\right) \cdot 11.10 /
$$

In other words, this subgroup satisfies the equation

$$
H_{\tau}(\varphi) \hat{S}_{\tau} H_{\tau}(\varphi)^{-1}=\hat{S}_{\tau}
$$

with $S_{\tau}$ given by /1.8/. The range of $\varphi=\varphi_{1}+i \varphi_{2}$ in this case is given by the inequalities

$$
-2 \pi(1+\tau)<\varphi_{1}<2 \pi(1+\tau), \quad-\infty<\varphi_{2}<\infty .
$$

In the next Section we proceed to an investigation of orbits generated by the above subgroup in the space of complex vectors. In particular, we are interested in the orbits as $\tau \rightarrow \infty$.
2. Orbits on the Complex Sphere

According to Eq. /1.9/ the final point of the vector $\mathrm{S}_{\tau}$ is situated on the complex sphere $\sum_{S / l+\tau}$ of radius $\frac{S}{1+\tau}$ which is non-zero for finite $\tau$ but tends to zero as $\tau^{+\infty}$. At any rate, ${\underset{\sim}{~}}_{\tau}$ has the little group as given by Eq. /1.10/. Let us fix the value of $\tau$ for the time being and see what a little group $H_{\tau}^{\prime}(\varphi)$ is obtained if another standard vector of the same length is chosen instead of $S_{T}$. The answer is trivial, since as a consequence of the homogeneity of $\Sigma_{S / l+\tau}$ there exists an AeSL $(2, C)$ which translates $S_{\tau}$ into $S_{\tau}^{\prime}$ :

$$
\hat{S}_{\tau}^{\prime}=A \hat{S}_{\tau} A^{-1}
$$

It follows then from Eq. /1.11/ thati

$$
H_{\tau}^{\prime}(\varphi) \hat{S}_{\tau}^{\prime} H_{\tau}^{\prime}(\varphi)^{-1}=\hat{S}_{\tau}^{\prime}
$$

where

$$
H_{\tau}^{\prime}(\varphi)=A H_{\tau}(\varphi) A^{-1}
$$

Thus the little group of an arbitrary complex vector $S_{\tau}^{\prime}(\varphi)$ is a group conjugate to $H_{\tau}(\varphi)$ given by Eq./l.10/, that is, $H_{\tau}^{\prime}(\varphi)$ is isomorphic to SO (2,C) when $\tau \rightarrow \infty$ and isomorphic to $T(2)$ as $\tau+\infty$.

Now, we are interested in orbits of a complex vector $\underset{\sim}{\underset{\sim}{\sim}}$ under the group $/ 2.3 /$. It is supposed that the final point of $\underset{\sim}{S}$ is situated on a complex sphere $\Sigma_{S}$ of non-zero radius/not to be confused with the sphere ${ }_{S} / 1+\tau /$. Under the group $H_{\tau}^{\prime}(\varphi)$ the vector describes the orbit

$$
\hat{S}_{\tau}(\varphi)=H_{\tau}^{\prime}(\varphi) \hat{\mathrm{S}} \mathrm{H}_{\tau}^{\prime}(\varphi)^{-1}
$$



Fig. 1

> Illustration of orbits on the complex sphere $\sum_{S}$. The orbit on the complex sphere $\Sigma$ S under the little group H, $(\varphi)$ is situated on the intersection of the complex sphere $\sum_{S}$ and a complex plane with a normal vector $\mathbb{N}_{\sim}$. In the limit $\tau \rightarrow \infty$ the orbit tends to the horosphere, which cannot be visualized so simply since in this case the normal becomes a complex vector of zero length (i.e. $\mathbb{N}_{\infty}^{2}=0$, though ${\underset{\sim}{N}}_{\infty} \neq 0)$.

Since $H^{\prime}(\varphi) \subset S L(2, C)$ the orbit is obviously situated on the sphere $\Sigma_{S}$. Moreover, it will be verified that the orbit lies in a complex plane ${\underset{\sim}{S}}_{\tau}(\varphi) \underset{\sim}{N_{\tau}}=C_{\tau}=$ const where ${\underset{\sim}{N}}_{\tau}$ proves to be indentical with $S_{\tau}^{\prime}$ given by Eq. /2.1/ Indeed, it follows from Eqs. /2.3/ and /2.4/ that

$$
\begin{align*}
& {\underset{\sim}{S}}_{\tau}(\varphi){\underset{\tau}{\tau}}_{\prime}^{\prime}=\frac{1}{2} \operatorname{Tr}\left(\hat{S}_{\tau}(\varphi) \hat{S}_{\tau}^{\prime}\right)= \\
& =\frac{1}{2} \operatorname{Tr}\left(H_{\tau}^{\prime}(\varphi) \hat{S}_{H_{\tau}^{\prime}}(\varphi)^{-1} \hat{S}_{\tau}^{\prime}\right)= \\
& =\frac{1}{2} \operatorname{Tr}\left(\hat{S}_{H_{\tau}^{\prime}}(\varphi)^{-1} \hat{S}_{\tau}^{\prime} H_{\tau}^{\prime}(\varphi)\right)= \\
& =\frac{1}{2} \operatorname{Tr}\left(\hat{S} \hat{S}_{\tau}^{\prime}\right)=\underset{\sim}{S}{\underset{S}{\tau}}_{\prime} \equiv C_{\tau}=\text { const. }
\end{align*}
$$

According to this equation one can associate with each orbit generated by the little group $H_{\tau}^{\prime}(\varphi)$ a normal vector ${\underset{\sim}{\sim}}_{\prime}^{\prime}$. The orbit can be given by the homogeneous coordinates ( ${\underset{\sim}{~}}_{\tau}^{\prime}, C_{\tau}$ ) ; nevertheless, apart from the singular case $C_{\tau}=0$, one can normalize $C_{\tau}$ tol by an appropriate dilatation of $S_{\tau}^{\prime}$. As the above statements are independent of the value of $\tau$, we can take the limit ${ }_{\tau \rightarrow \infty}$, which produces horospheres. So, according to /2.3/ and /2.4/ the horospheres on $\sum_{\mathrm{s}}$ are orbits described by the horospheric subgroup

$$
\left.H_{\infty}^{\prime}(\varphi)=\left(\begin{array}{ll}
\alpha & \beta \\
\gamma & \delta
\end{array}\right)\left(\begin{array}{cc}
1 & -i \varphi \\
0 & 1
\end{array}\right)\left(\begin{array}{ll}
\alpha & \beta \\
\gamma & \delta
\end{array}\right)^{-1}=A H_{\infty} A^{-1} \quad \right\rvert\, \alpha \delta-\gamma \beta=1 /
$$

for fixed A. Taking into account Eq. /1.9/ and the fact that transformation /2.1/ leaves the longth of ${\underset{\sim}{\sim}}_{\tau}$ unchanged, we get for $\tau \rightarrow \infty$

$$
\left(\mathrm{s}_{\infty}^{\prime}\right)_{1}^{2}+\left(\mathrm{s}_{\infty}^{\prime}\right)_{2}^{2}+\left(\mathrm{s}_{\infty}^{\prime}\right)_{3}^{2}=0
$$

Thus, when the $\mathrm{so}(2, \mathrm{c})$ group deforms into the horospheric group as $\tau \rightarrow \infty$, the normal vector $\underset{\sim}{N}{ }_{\tau} \equiv \underset{\sim}{S}{ }_{\tau}^{\prime}$ characteristic for the orbits arrives at the complex sphere of zero radius, that is ${\underset{N}{N}}_{2}^{2}={\underset{\sim}{\infty}}_{\prime}^{\prime 2}=0 \quad$. It is concluded that horospheres of $\Sigma$ are determined by the equation ${\underset{\sim}{N}}_{\infty}=1$, where $N_{\infty}^{2}=0$.

[^1]At this point a remark is in order. Namely, we did not investigate the question whether in Eq. /2.5/ the normal vector of the plane of horospheres is unique up to a factor. From a more detailed investigation which for the sake of brevity is left to the reader, the following can be shown. A single fixed point $S$ of the space $\Sigma_{S}$ is crossed by a one-parametric manifold of horospheres. These are second order curves which, generally speaking, determine unambiguously a plane with a normal of zero length, as indicated above. However, in the manifold of horospheres crossing a fixed point there are two positions where the horosphere degenerates into a complex straight line. These lines can be given in the form

$$
{\underset{\sim}{\infty}}_{\infty}(\varphi)=\underset{\sim}{A} \varphi+\underset{\sim}{S}, \quad \underset{\sim}{S}(\varphi)=\underset{\sim}{B} \varphi+\underset{\sim}{S}
$$

with

$$
\begin{aligned}
& \underset{\sim}{A}=\left(A_{1}+i A_{2}, A_{1}-i A_{2}, A_{3}\right)=\left(S \frac{S_{+}^{2}}{\left(S+S_{3}\right)^{2}},-S, S \frac{S_{+}}{S+S_{3}}\right) \\
& \left.\underset{\sim}{B}=\left(B_{1}+i B_{2}, B_{1}-i B_{2}, B_{3}\right)=\left(-S \frac{S_{+}^{2}}{\left(S-S_{3}\right)^{2}}, S, S \frac{S_{+}}{S-S_{3}}\right) \quad \right\rvert\, S \pm S_{3} \neq 0 /
\end{aligned}
$$

where the usual $S_{ \pm}=S_{1} \pm i S_{2}$ notation is used. Therefore, each point $\underset{\sim}{S}$ is crossed by two straight horospheres, that are determined by the position of
$\underset{\sim}{S}$ alone. These horospheres can be called horospheres of the second kind, as distinguished from those of the tirst kind, which are in one-to-one correspondence with the vectors of the sphere of zero radius. An analogous situation is encountered in the familiar case of the three-dimensional real one-sheeted hyperboloid [I].

## 3. Deformation of Unitary Representations

To demonstrate the deformation on unitary representations let us consider the linear fractional mapping of the $z-p l a n e$ which is a factor space $\operatorname{SL}(2, C) /\left(\begin{array}{ll}\alpha & \beta \\ 0 & \alpha^{-1}\end{array}\right):$

$$
z^{\prime}=\frac{\alpha z+\beta}{\gamma z+\delta}
$$

In the case of the $S O(2, c)$ subgroup given by $/ 1.5 /$ this reduces to

$$
z^{\prime}=e^{-i \varphi} z=e^{\varphi_{2}} e^{-i \varphi} 1
$$

which is a rotation followed by a dilatation. In a similar way, the horospheric transformation on the $z$-plane takes the form of an Euclidean displacement

$$
z^{\prime}=z-i \varphi=z+\left(\varphi_{2}-i \varphi_{1}\right)
$$

The interpolating subgroup given by Eq. /l.l0/ accomplishes a transformation similar to that of $/ 3.1 /$ on the displaced $z$-plane, i.e.

$$
z^{\prime}+\tau=e^{-i \frac{\varphi}{1+\tau}}(z+\tau)
$$

For realizing unitary representations the representation on the familiar
$\varphi(z)$ functions will be used [1]. Action of an element $A=\left(\begin{array}{ll}\alpha & \beta \\ \gamma & \delta\end{array}\right)$ on these functions is defined as

$$
\mathrm{T}_{\mathrm{A}} \varphi(\mathrm{z})=(-\gamma z+\alpha)^{2 j}\left(-\gamma^{*} z^{*}+\alpha^{*}\right)^{2 k} \varphi\left(\frac{\delta z-\beta}{-\gamma z+\alpha}\right)
$$

where

$$
j=\frac{1}{2}\left(j_{\circ}-1+i \sigma\right), \quad k=\frac{1}{2}\left(-j_{\circ}-1+i \sigma^{*}\right)
$$

Here $j_{o}$ takes integer and half-integer values, while $\sigma$ is an arbitrary complex number. In what follows we restrict ourselves to the principal series of unitary representations for which $\sigma$ is real. In Eq. /3.2/ representations are defined by displacement from the left which results in the following form of infinitesimal generators:

$$
\begin{align*}
& J_{+}=J_{1}+i J_{2}=-\frac{\partial}{\partial z}, \quad K_{+}=K_{1}+i K_{2}=2 k z^{*}-z^{* 2} \frac{\partial}{\partial z^{*}} \\
& J_{-}=J_{1}-i J_{2}=-2 j z+z^{2} \frac{\partial}{\partial z}, K_{-}=K_{1}-i K_{2}=\frac{\partial}{\partial z^{*}} \\
& J_{3}=j-z \frac{\partial}{\partial z}, \\
& K_{3}=-k+z^{*} \frac{\partial}{\partial z^{*}}
\end{align*}
$$

These generators are related to the generators of spatial rotations about $k^{\text {th }}$-axis $M_{k} / k=1,2,3 /$ and to the generators of boosts along $k^{\text {th }}$-axis $N_{k}$ as $J_{k}=\frac{1}{2}\left(M_{k}+i N_{k}\right), \quad K_{k}=\frac{1}{2}\left(M_{k}-i N_{k}\right)$.

Spherical functions in $S O(2, C)$ basis satisfy the eigenvalue equations

$$
\begin{align*}
& J_{3} \varphi_{\mathrm{mm}^{*}}^{\circ}\left(\mathrm{z}, \mathrm{z}^{*}\right)=\mathrm{m} \varphi_{\mathrm{mm}}^{\circ} \mathrm{O}\left(\mathrm{z}, \mathrm{z}^{*}\right) \\
& \mathrm{K}_{3} \varphi_{\mathrm{mm}}{ }^{\circ}\left(\mathrm{z}, \mathrm{z}^{*}\right)-=\mathrm{m}^{*} \varphi_{\mathrm{mm}}^{0}\left(\mathrm{z}, \mathrm{z}^{*}\right)
\end{align*}
$$

where

$$
m=\frac{1}{2}(\mu+i v), \quad m^{*}=\frac{1}{2}(\mu-i v)
$$

with $\mu=0, \pm \frac{1}{2}, \pm 1, \ldots$ and $v$ continuous. The above basis is a generalization of finite dimensional spinors to the unitary case where $m$ and $m^{*}$ correspond to undotted and dotted indices of spinors. Unitary spinors can be succesfully applied to the evaluation of matrix elements of unitary representations of the Lorentz group, namely, they simplify to considerable extent the results obtained in angular momentum basis. $[7,8,9,10]$.

The solution of $/ 3.5 /$ is

$$
\varphi_{m m^{*}}^{0}\left(z, z^{*}\right)=\frac{1}{2 \pi} z^{j-m} z^{* k+m^{*}}
$$

13.71

The requirement of single-valuedness on the complex $z-p l a n e$ yields the condition $2 \mu=$ integer or, to be strict, $\mu$ takes integer and half-integer values along with $j_{0}$. The functions /3.7/ are normalized as

$$
\left\langle m^{\prime} m^{\prime *} \mid m m^{*}\right\rangle=\frac{i}{2} \int d z d z^{*} \varphi_{m}^{\prime} m^{\prime *}\left(z z^{*}\right)^{*} \varphi_{m m^{*}}^{o}\left(z, z^{*}\right)=\delta_{\mu^{\prime} \mu} \delta\left(v^{\prime}-v\right)
$$

The horospheric group as given by Eq. /l.7/ is generated by the Hermitean generators $M_{1}-N_{2}$ and $M_{2}+N_{1}$ or - equivalently - by the non--Hermitean generators $J_{+}$and $K_{\text {_ }}$. In this basis spherical functions are solutions of the eigenvalue equations

$$
\begin{align*}
& J_{+} \varphi_{m m^{*}}^{\infty}\left(z, z^{*}\right)=m \varphi_{m m^{*}}^{\infty}\left(z, z^{*}\right), \\
& K_{-} \varphi_{m m^{*}}^{\infty}\left(z, z^{*}\right)=m^{*} \varphi_{m m^{*}}^{\infty}\left(z, z^{*}\right)
\end{align*}
$$

## References

[1] I.M. Gelfand, M.I. Graev and N.Ya. Vilenkin Generalized Functions, Vol.5. Academic Press, New York and London, 1966.
[2] E. Inonu and E.P. Wigner, Proc. Natl. Acad. Sci. /U.S./ 32, $510 / 1953 /$.
[3] R. Hermann, Fourier Analysis on Groups and Partial Wave Analysis. W.A. Benjamin, Inc.New York, 1969, p. 35-47.
[4] K. Szegő and K. Tóth, Journ.Math. Phys. 12, 486 and 853 /1971/。
[5] A.J. Macfarlane, Journ.Math.Phys. 3, 1116 /1962/.
[6] M. Huszár, Commun.Math.Phys, 23, 132 /1971/.
[7] H. Joos and R. Schrader, DESY Preprint 68/40, Hamburg, 1968.
[8] M.Huszár and J. Smorodinsky, JINR Preprint E2-4225 /1968/.
[9] M. Carmeli, Journ, Math. Phys. 11, 1971 /1970/.
[10] M. Carmeli and S. Malin, Journ. Math. Phys. 12, 225 /1971/.

$$
61.973
$$



$\qquad$




$\qquad$

$\qquad$
 $\qquad$


Kiadja a Központi Fizikai Kutató Intézet
Felelős kiadó: Kiss Dezső, a KFKI Nagyenergiáju Fizikai Tudományos Tanácsának elnöke
Szakmai lektor : Sebestyén Åkos és Szegő Károly Nyelvi lektor: T. Wilkinson
Példányszám: 330 Törzsszám : 72-7069
Készült a KFKI sokszorositó üzemében, Budapest 1972. julius hóban


[^0]:    */Under proper Lorentz transformations, a complex vector $S$ transforms like B+iE , where $\underset{\sim}{B}$ and $\underset{\sim}{E}$ are the magnetic and electric field strengths.
     electrodynamics as well.

[^1]:    */
    The real and jomaginary parts of a complex vector on the complex sphere of zero radius are quantities analogous to the field strengths $\underset{\sim}{B}$ and $E$ of an olectromagnotic plane wave, where $\underset{B_{2}}{B}$ and $\underset{\underset{E}{E}}{ }$ are of the same modulus and are perpendicular to each other: ${\underset{N}{2}}^{2}-{\underset{E}{2}}^{2}=0, \underset{\sim}{\mathrm{BE}}=0$.

